Phenomenology of primordial black hole from a first order phase transition

Speaker: Po-Yen Tseng

(Department of Physics, NTHU)

References: 2305.14399, 2304.10084, 2212.13035, 2209.01552





OMU-NTHU Meeting, Feb 6-9 2025

Goal for particle physics

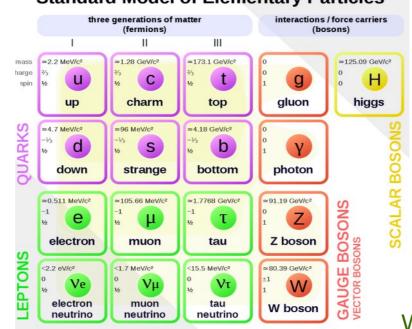
The periodic table:

Period v																			Noble gases
														Some	eleme	nts nea	ar		<u> </u>
N	1													the da	shed s	staircas	e are		2
Nonmetals	H			sometimes called <i>metal</i>									netallo	ids	He				
Metals 2	3	4												5	6	7	8	9	10
-	Li	Be												В	С	Ν	0	F	Ne
3	11	12		Transition metals								13	14	15	16	17	18		
0	Na	Mg		(sometimes excluding group 12) AI									Si	Р	S	CI	Ar		
4	19	20		21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36
	к	Ca		Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
5	37	38		39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54
0	Rb	Sr		Y	Zr	Nb	Мо	Тс	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	1	Xe
6	55	56	La to Yb	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86
0	Cs	Ba		Lu	Hf	Та	W	Re	Os	Ir	Pt	Au	Hg	TI	Pb	Bi	Po	At	Rn
7	87	88	Ac to No	103	104	105	106	107	108	109	110	111	112	113	114	115	116	117	118
,	Fr	Ra	/ 10 10 110	Lr	Rf	Db	Sg	Bh	Hs	Mt	Ds	Rg	Cn	Nh	FI	Mc	Lv	Ts	Og
	s-hl	ock	f-block	ľ.				d-bl	ock						n-hlo	ck (ex	luding	1 He)	
	s-block f-block d-block (plus He)									p-block (excluding He)									
	(prac	, 110)																	
Lanthanides				57	58	59 Dr	60	61	62	63 5	64	65	66	67	68	69 T	70 X/b		
				La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Но	Er	Tm	Yb		
Actinides			89	90	91 D	92	93	94	95	96	97	98	99	100	101	102			
				Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No		

Wikipedia: Periodic table

OMU-NTHU meeting 2025

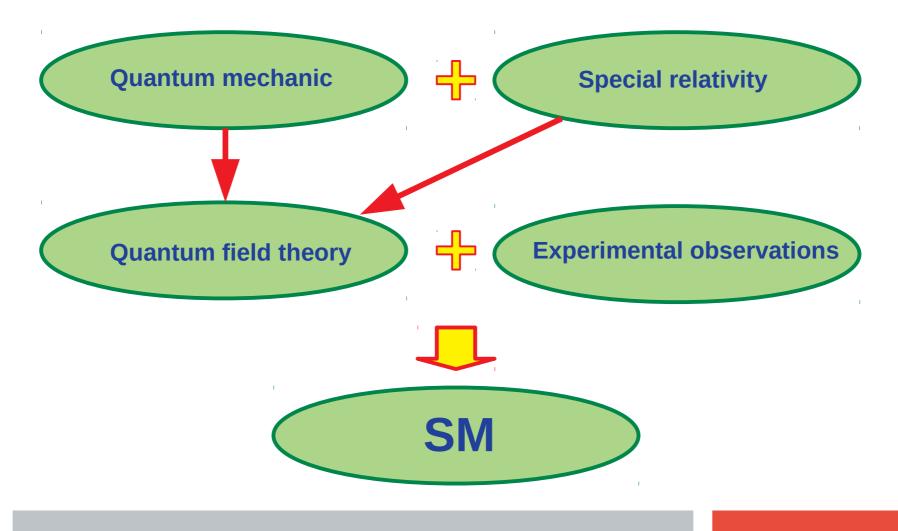
 (1978-present): SM unifies three of the four fundamental forces (electromagnetic, weak, and strong interactions), and classifies all known elementary particles.



Standard Model of Elementary Particles

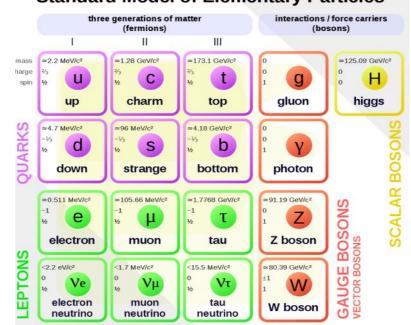
Wikipedia: standard model

OMU-NTHU meeting 2025



OMU-NTHU meeting 2025

 (1978-present): SM unifies three of the four fundamental forces (electromagnetic, weak, and strong interactions), and classifies all known elementary particles.

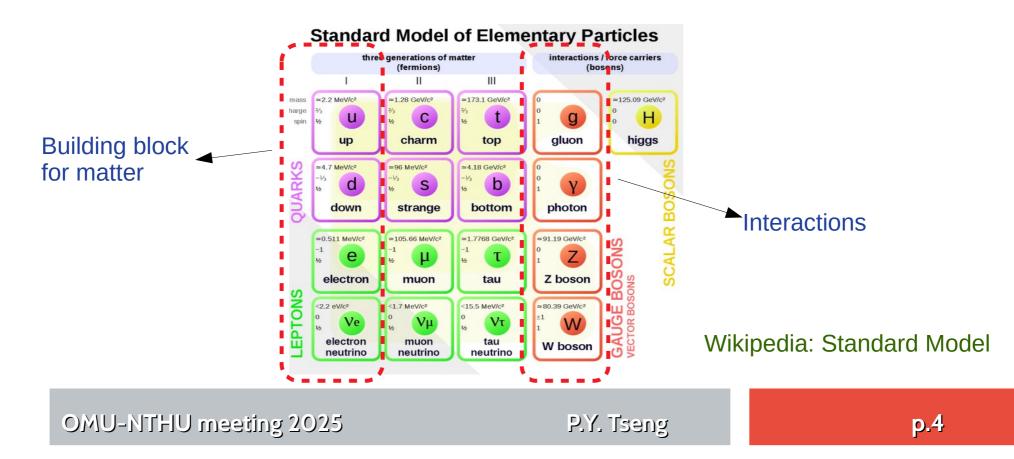


Standard Model of Elementary Particles

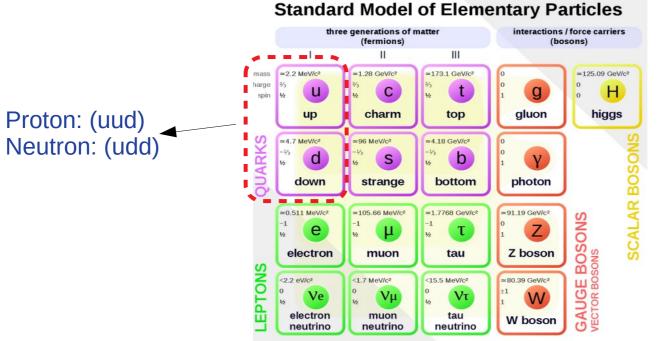
Wikipedia: Standard Model

OMU-NTHU meeting 2025

 (1978-present): SM unifies three of the four fundamental forces (electromagnetic, weak, and strong interactions), and classifies all known elementary particles.



 (1978-present): SM unifies three of the four fundamental forces (electromagnetic, weak, and strong interactions), and classifies all known elementary particles.

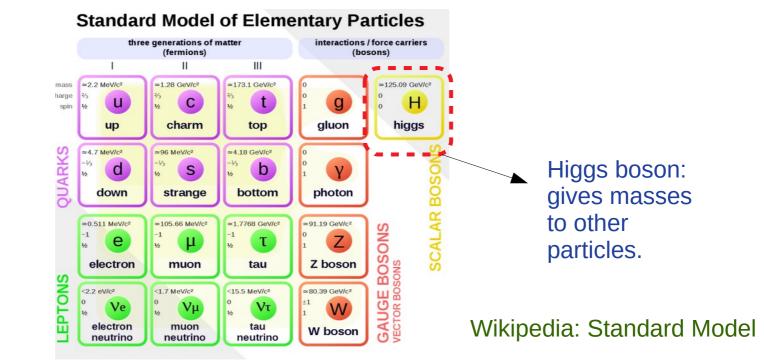


Wikipedia: Standard Model

OMU-NTHU meeting 2025

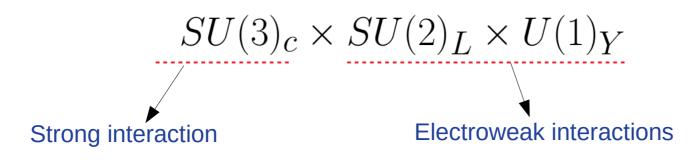


 (1978-present): SM unifies three of the four fundamental forces (electromagnetic, weak, and strong interactions), and classifies all known elementary particles.



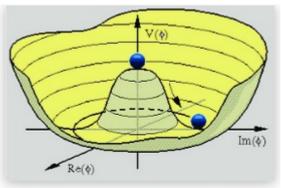
OMU-NTHU meeting 2025

- (1978-present): SM unifies three of the four fundamental forces (electromagnetic, weak, and strong interactions), and classifies all known elementary particles.
- Lagrangian of SM is invariant under the symmetry transformation:



Standard Model: Higgs mechanism

Potential of Higgs field:



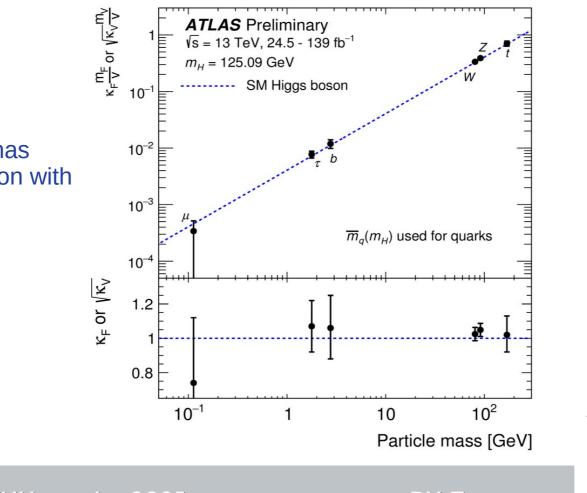
Explanation of the Brout-Englert-Higgs (BEH) ... researchgate.net

$$V(H) = -\mu^2 |H|^2 + \lambda |H|^4$$

• Minimum of potential: $H \neq 0$

 Higgs mechanism prediction that particle's mass proportional to its coupling strength with Higgs.

Heavier particle has stronger interaction with Higgs boson.



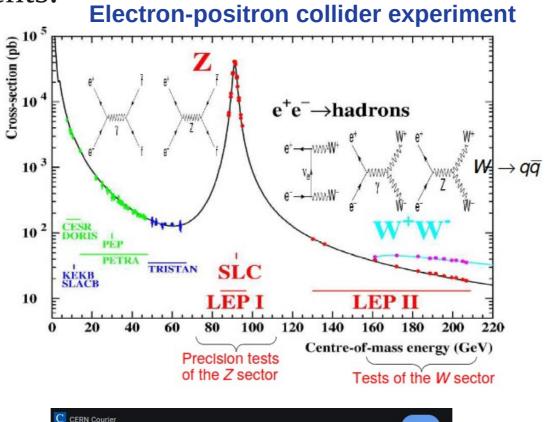
ATLAS Higgs physics group

OMU-NTHU meeting 2025

Success of SM

 Successful theory should be tested by many observations and experiments.

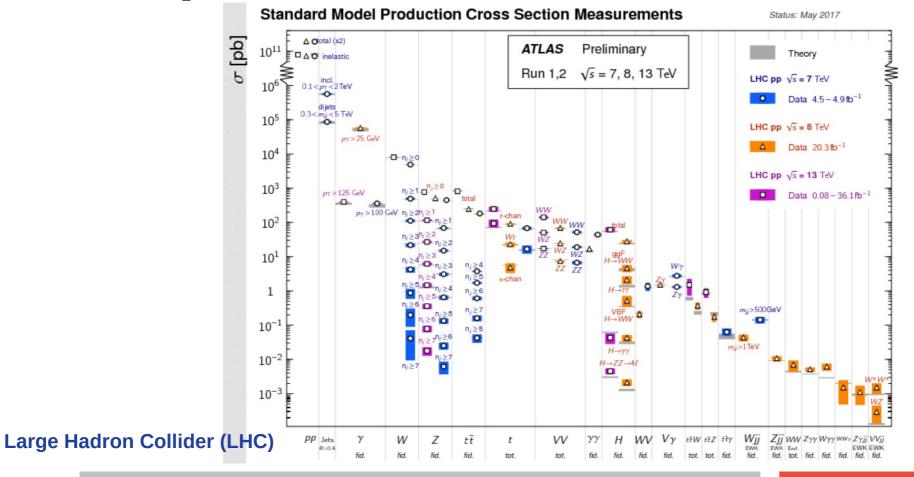
Revisiting the b revolution – CERN Courier



OMU-NTHU meeting 2025

Success of SM

 Successful theory should be tested by many observations and experiments.
 Kretzschmar, Jan, ATLAS, CMS, CERN Document: arXiv:1803.10800

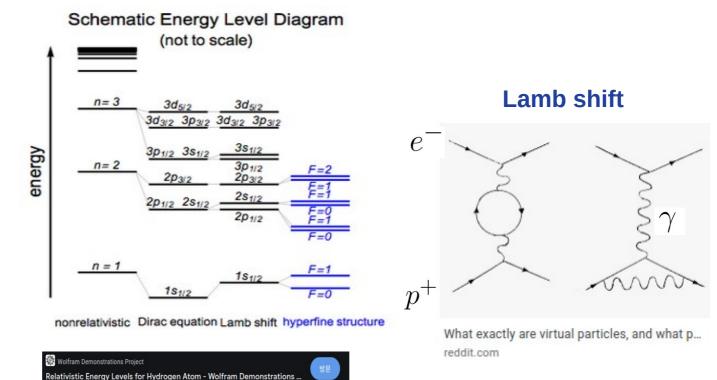


OMU-NTHU meeting 2025

Success of SM

 Successful theory should be tested by many observations and experiments.

Energy spectra of Hydrogen atom:



P.Y. Tseng

OMU-NTHU meeting 2025

p.10

Members in our group

Postdoc:



Jan Tristram Acuna

PhD:



Yu-Min Yeh



Kuan-Yen Chou

Master:





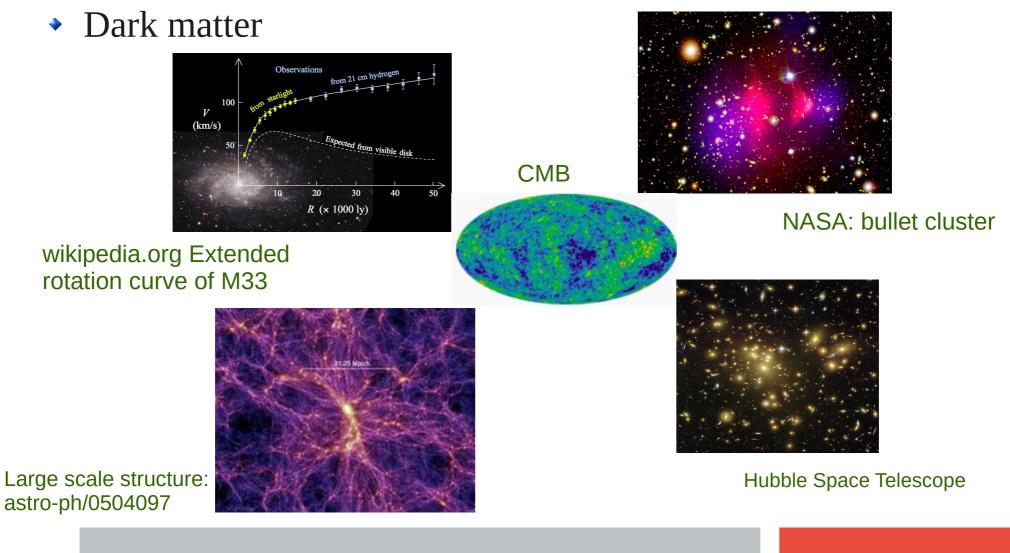
Leah, Ya-Ling Lin Chang-Yu Dai

OMU-NTHU meeting 2025



Dark Matter (DM):

Indirect evidences of DM



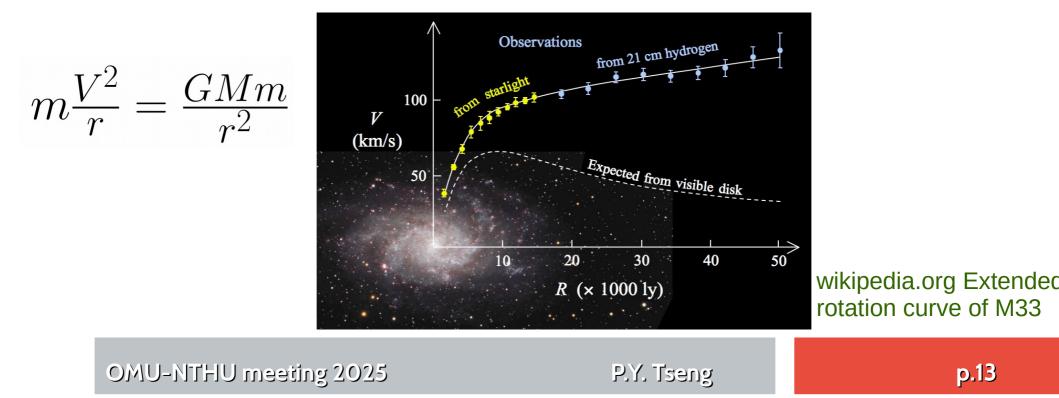
OMU-NTHU meeting 2025

P.Y. Tseng

p.12

Indirect evidences of DM

- Rotational curve of spiral galaxy: gravitational pulling force equals to centrifugal force.
- However, the observed rotational velocity is much faster than expected from visible disk.



Indirect evidences of DM

- We believe that the galaxy disk is surrounded by spherical DM halo.
- Total DM mass is 5 times heavier than galaxy disk.

 Image: Second state of the subhalos of the subh

P.Y. Tseng

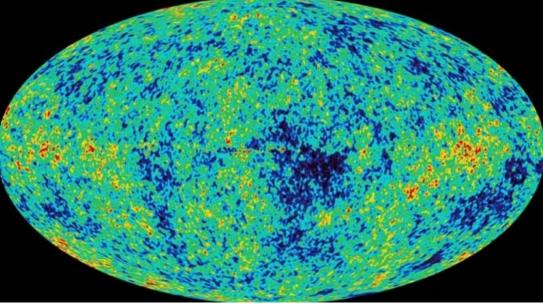
Dark matter do not interact with photon.

OMU-NTHU meeting 2025

p.14

Indirect evidences of DM: CMB

 Observation of Cosmic Microwave Background (CMB) by satellite Planck

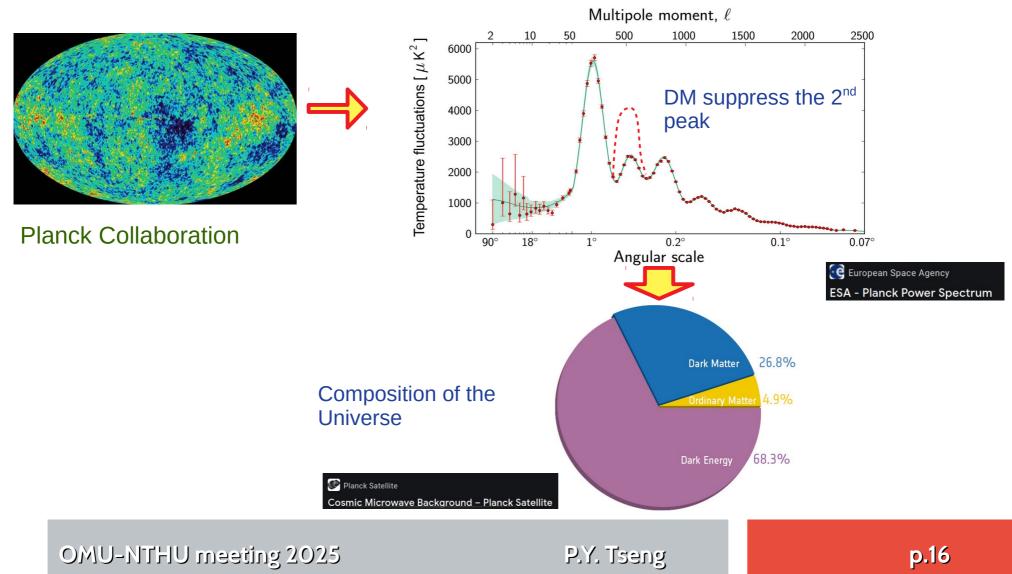


Planck Collaboration

T=2.725 K. Temperature is very uniform universe. The color indicate the temperature fluctuation by 1/10000

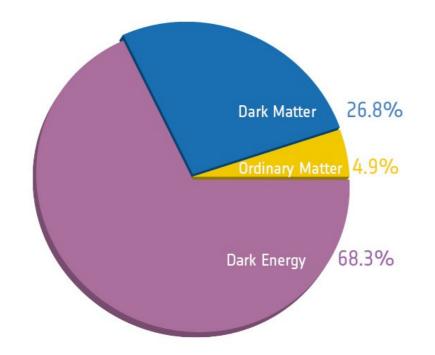
Indirect evidences of DM: CMB

• Expand the temperature by spherical harmonic functions:



DM relic abundance

 Dark matter relic abundance is about 25% of our Universe. Ordinary matter(SM particles, baryon, matter interact with photon) is about 5%.

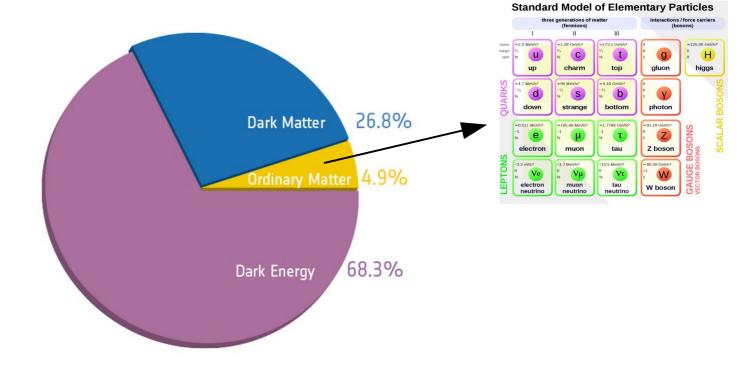




OMU-NTHU meeting 2025

DM relic abundance

 Dark matter relic abundance is about 25% of our Universe. Ordinary matter(SM particles, baryon, matter interact with photon) is about 5%.





OMU-NTHU meeting 2025

p.17

DM candidates

- Properties of **Dark matter:**
- I).Charge neutral.
- II).Lifetime longer than age of Universe.
- III).Non-relativistic.

(bosons) L Ш Ш mass 2.2 MeV/c2 ≃1.28 GeV/c2 ≃173.1 GeV/c2 ≈125.09 GeV/c² harge ⅔ С g u spin charm top gluon higgs up QUARKS ≃4.7 MeV/c ≈96 MeV/c² ≃4.18 GeV/c³ -1/2 S b d Y bottom down strange photon ≃0.511 MeV/c2 ≈105.66 MeV/c ≈1.7768 GeV/c³ ≈91.19 GeV/c2 BOSONS Ζ е μ τ muon Z boson electron tau EPTONS. **GAUGE** I <2.2 eV/c2 <1.7 MeV/c² <15.5 MeV/c² ≈80.39 GeV/c2 Ve νμ ντ W tau electron muon W boson neutrino neutrino neutrino

three generations of matter

(fermions)

Standard Model of Elementary Particles

Neutron decays. Neutrinos are relativistic.

OMU-NTHU meeting 2025

P.Y. Tseng

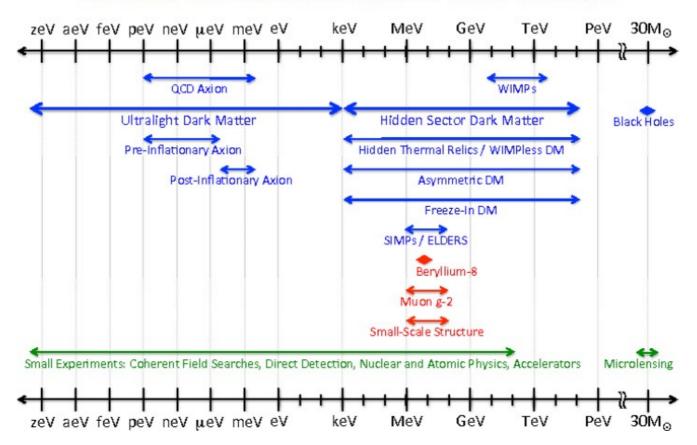
interactions / force carriers

н

SCALAR BO

DM candidates

• DM mass:



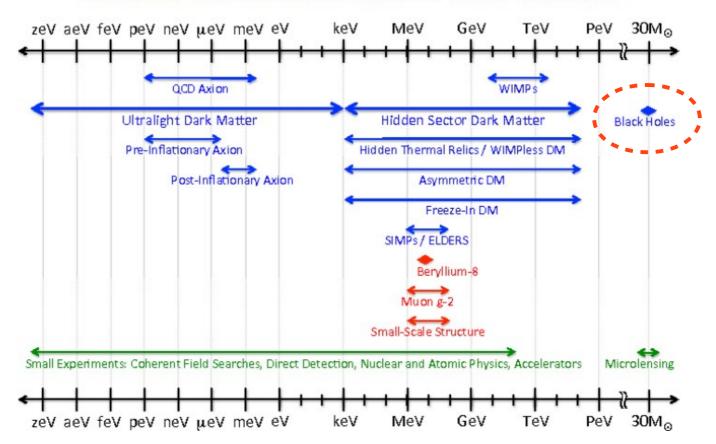
Dark Sector Candidates, Anomalies, and Search Techniques

CERN document Server: US:Cosmic Visions: New ideas in dark matter 2017

OMU-NTHU meeting 2025

DM candidates

DM mass:



Dark Sector Candidates, Anomalies, and Search Techniques

CERN document Server: US:Cosmic Visions: New ideas in dark matter 2017

OMU-NTHU meeting 2025

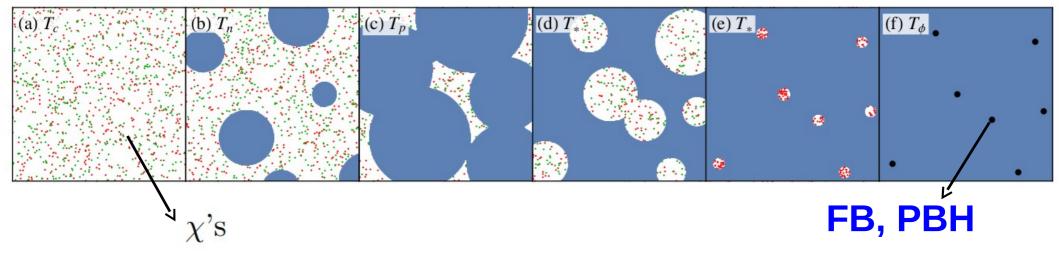
Research topics:

First order phase transition and macroscopic DM

PBHs from FOPT

• Schematic for **PBHs** formation form **FOPT**:





The χ's trapped outside the bubble form macroscopic DM (Fermi Ball or PBH). Consequently, the mass of PBH can be much lighter than the Sun.



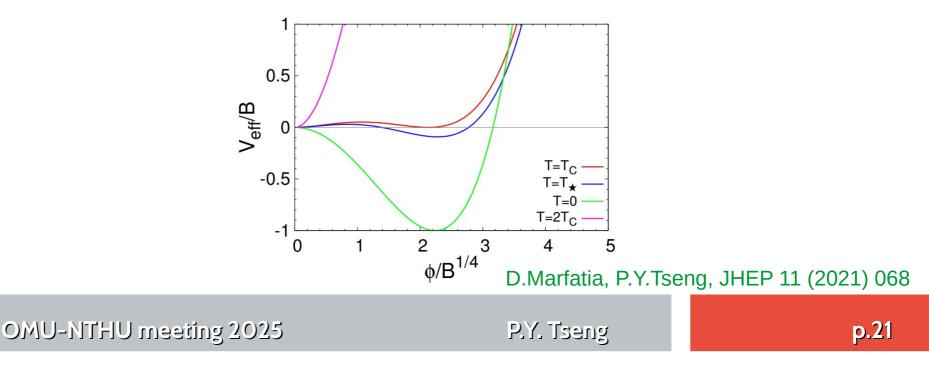
Model framwork

 The origin of DM mass may come from the spontaneous symmetry breaking inducing by another scalar.

$$\mathcal{L} \supset \bar{\chi} i \partial \!\!\!/ \chi - g_{\chi} \phi \bar{\chi} \chi - V_{\text{eff}}(\phi, T)$$

$$m_{\chi} \simeq g_{\chi} \langle \phi \rangle$$

• We consider **1**st order phase transition (FOPT).



Light PBHs from FOPT

• PBHs with 1E-17 solar mass are produced from MeV energy scale **FOPT**.

	BP-1	BP-2	BP-3	BP-4	BP-5	BP-6
λ	0.061	0.110	0.195	0.087	0.150	0.158
$B^{1/4}/{ m MeV}$	75.14	13.81	1.501	1.261	0.121	2.999
$C/{ m MeV}$	0.249	0.462	0.078	0.052	0.011	0.325
D	0.596	1.458	1.119	0.596	1.418	0.519
g_{χ}	1.088	1.301	1.011	1.289	0.983	1.228
η_{χ}	1.03×10^{-9}	1.28×10^{-10}	1.64×10^{-12}	1.21×10^{-15}	2.59×10^{-18}	6.26×10^{-17}
m/MeV		0.120	0_259	0.394	0-341	1.704
$T_{ m SM\star}/ m MeV$	94.68	14.63	0.895	2.104	0.164	4.774
$T_{\star}/{ m MeV}$	53.16	6.143	0.421	0.868	0.052	2.287
$T_f/{ m MeV}$	59.63	6.888	0.472	1.023	0.068	2.571
$T_{\phi}/{ m MeV}$	53.09	6.045	0.415	0.857	0.050	1.950
$S_3(T_\star)/T_\star$	155		166	171	180	
$M_{ m PBH}/M_{\odot}$	2.92×10^{-16}		1.19×10^{-17}	1.93×10^{-18}	3.91×10^{-19}	
$Q_{ m FB}$	1.26×10^{42}		$5.96 imes10^{42}$	5.01×10^{41}	$7.58 imes 10^{41}$	
eta^\prime	$2.80 imes 10^{-17}$	2.54×10^{-19}	$7.78 imes10^{-23}$	4.45×10^{-26}	$5.75 imes10^{-30}$	8.97×10^{-28}
α	1.48×10^{-2}	$7.40 imes 10^{-3}$	1.20×10^{-2}	1.12×10^{-2}	$1.35 imes 10^{-2}$	$1.30 imes 10^{-2}$
eta/H_{\star}	4.41×10^{3}	$9.36 imes 10^3$	3.21×10^4	3.25×10^3	4.94×10^3	$2.64 imes 10^3$
v_w	0.904	0.904	0.904	0.930	0.963	0.905
$v_{\phi}(T_{\star})/\mathrm{MeV}$	224	23.1	1.426	3.821	0.247	8.157
$dM_{ m FB}/dQ_{ m FB}/{ m MeV}$	258	28.3	1.980	4.264	0.573	10.89
$\Omega_{\rm PBH}h^2$	0.079	1.12×10^{-3}	1.09×10^{-6}	1.52×10^{-9}	2.15×10^{-13}	6.35×10^{-29}
$\Delta N_{ m eff}$	0.218	0.126	0.208	0.146	0.147	0.221

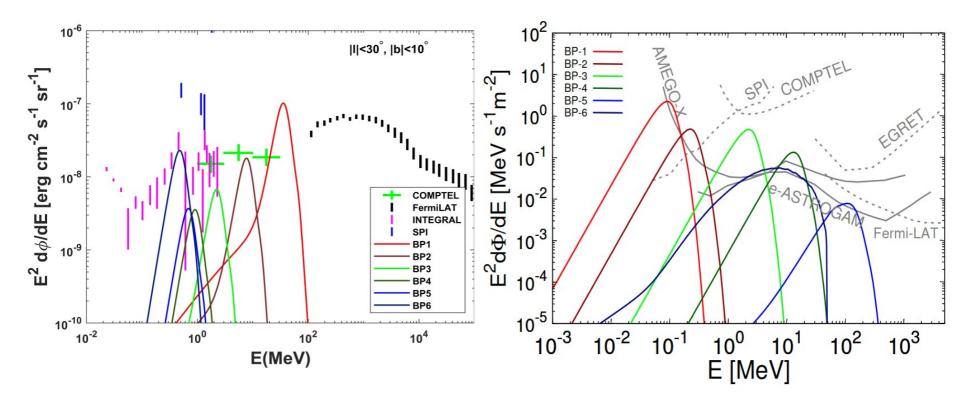
D.Marfatia, P.Y.Tseng, JHEP 08 (2022) 001

OMU-NTHU meeting 2025



Gamma-ray from PBH evaporation

• Galactic/Extragalactic gamma-ray spectra and GC 511 keV line from PBH evaporation:



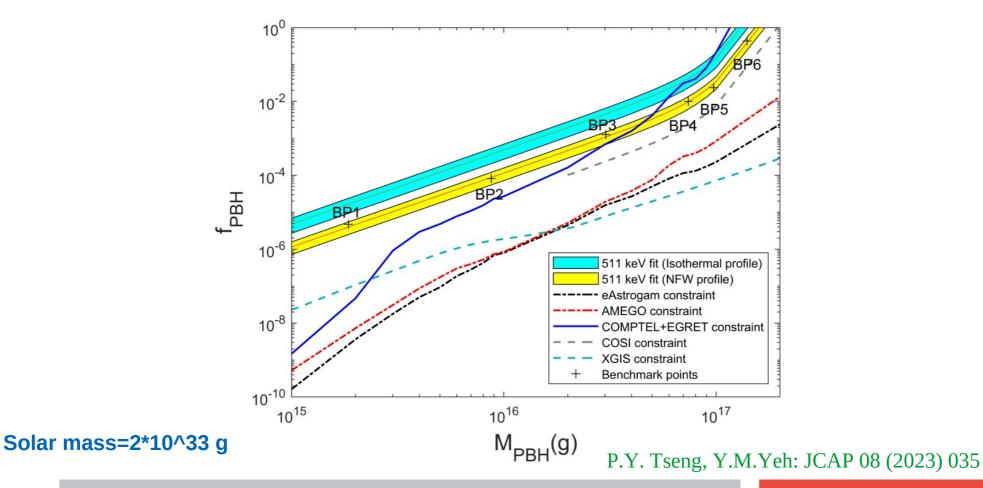
P.Y. Tseng, Y.M.Yeh: JCAP 08 (2023) 035

OMU-NTHU meeting 2025

p.24

GC 511 keV gamma-ray and PBH

Produced black hole mass can be much lighter than our Sun:

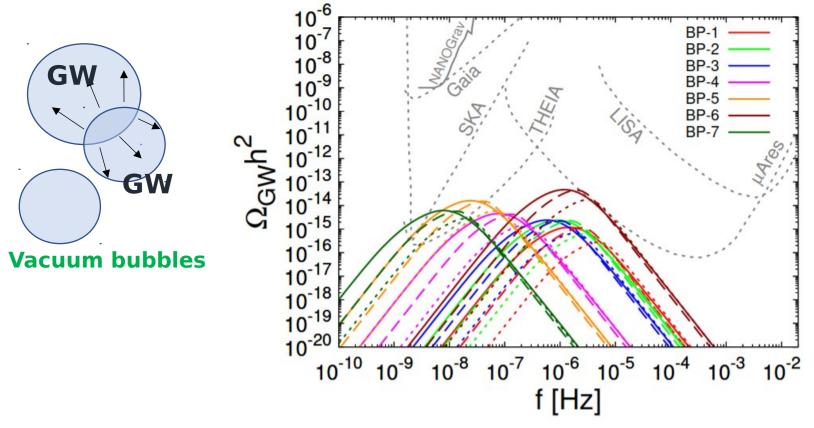


OMU-NTHU meeting 2025



Gravitational wave signals

• GW spectra from the FOPT with bubble size distribution:



D. Marfatia, P.Y. Tseng, Y.M.Yeh: 2407.15419



Axion star and Fast Radio Bursts

Introduction

- **QCD axion**: solve strong CP problem.
- Axion-like particle (ALP): is a psuedoscalar boson, its mass is not linear proportional to the couplings to SM particles.
- **ALP** remains one of the dark matter candidates.
- **ALP** can couple to photons: $g_{a\gamma\gamma}aF^{\mu\nu}\tilde{F}_{\mu\nu}$

Introduction

- ALPs are produced at high occupancy in the early Universe by a misalignment mechanism.
- ALPs form the Bose-Einstein condensate (BEC) if they are in thermal equilibrium.
- ALPs BEC is driven by attractive gravitational interaction, result in gravitational bound object *axion stars* (or *axion clumps*, *boson stars*).

Profile of Axion star

- Hamiltonian carries a global U(1) symmetry, which implies the total number of axion is conserved: $N \equiv \int d^3x |\psi|^2$
- We adopt the Gaussian ansatz for spatial wavefunction:

$$\psi_{\rm G}(r;R) = \sqrt{\frac{N}{\pi^{3/2}R^3}} \exp\left(-\frac{r^2}{2R^2}\right), \quad f_{\rm G}(x) = \exp\left(-x^2/2\right).$$

Minimizing the Hamiltonian:

$$N = 2.4258 \times 10^{60} \times \alpha \left(\frac{10^{-5} \text{ eV}}{m_a}\right)^2 \left(\frac{f_a}{10^{12} \text{ GeV}}\right) \left(\frac{0.3}{|\gamma|}\right)^{\frac{1}{2}}$$

$$R = 2.2830 \times 10^4 \text{ m} \times \left(\frac{1 + \sqrt{1 \mp \alpha^2}}{\alpha}\right) \left(\frac{10^{-5} \text{ eV}}{m_a}\right) \left(\frac{10^{12} \text{ GeV}}{f_a}\right) \left(\frac{|\gamma|}{0.3}\right)^{\frac{1}{2}}$$

$$M_{\text{AS}} = 2.1716 \times 10^{-11} M_{\odot} \times \alpha \left(\frac{10^{-5} \text{ eV}}{m_a}\right) \left(\frac{f_a}{10^{12} \text{ GeV}}\right) \left(\frac{0.3}{|\gamma|}\right)^{\frac{1}{2}}.$$

Phys.Rev.D 104 (2021) 12, 123012. [2109.04283]

OMU-NTHU meeting 2025

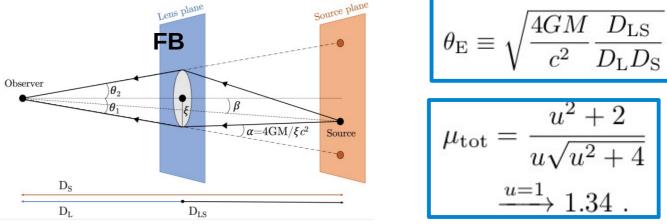
P.Y. Tseng



 f^2

Microlensing

These FB mass and radius ranges can induce microlensing effects.



D.Croon, D. McKeen, N. Raj: 2002.08962

 The separating angle of two images of the background star are too small to be resolved, but we can observe the sudden luminosity enhancement of the star.

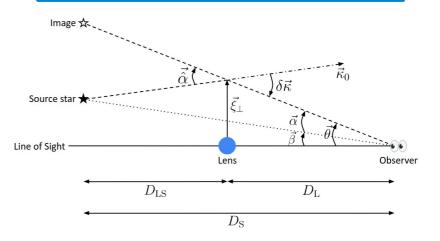
P.Y. Tseng

Light propagation in axion medium

- Deflection angle increases in longer wavelength, i.e lower frequency.
 J.T.Acuna, K.Y.Chou, P.Y.Tseng, arXiv:2501.07176
- The change in momentum of propagating photon

$$\left(\frac{\delta\vec{\kappa}}{\kappa_0}\right)_{\rm axion} \approx \frac{g_{a\gamma\gamma}^2}{8\kappa_0^2} \int_{-\infty}^{\infty} dt' \; \vec{\nabla}_{\perp} \left(\partial a\right)^2$$

$$\left(\frac{\delta\vec{\kappa}}{\kappa_0}\right)_{\text{gravity}} = -2\int_{-\infty}^{\infty} dz \; \vec{\nabla}_{\perp} \Phi$$

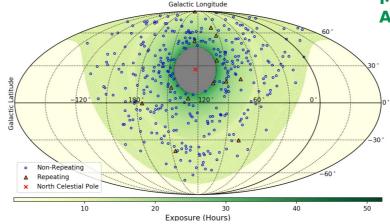


P.Y. Tseng

OMU-NTHU meeting 2025

Fast Radio Bursts (FRBs)

 FRBs are short-duration radio transients emitted by extragalactic sources.



Mandana Amiri et al. CHIME/FRB, Astrophys. J. Supp., 257(2):59, 2021.

Figure 1.7: Sky map of FRBs detected by CHIME/FRB [34].

- Advantages from FRBs: I) observed in radio frequency, II) short-pulse signature.
- Lens effect increases in low frequency.
- The observation can resolve two images due to time delay.

Fast Radio Bursts (FRBs)

 Signal requirements: I) Magnification ratio between two images < 5, ii) Time difference between two images is longer than microsecond. Julian B. Munoz et al. Phys.Rev.Lett 117 (2016) 9,

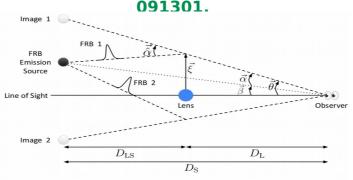


Figure 4.1: Geometrical setup of the FRB lensing event.

K.Y.Chou's thesis

Signal criteria:

$$\Delta t = |t(\theta_+) - t(\theta_-)|$$
$$R_f \equiv \frac{|\mu(\theta_+)|}{|\mu(\theta_-)|}$$

$$\Delta \bar{t} = 1 \,\mu \mathrm{s}$$

 $R_{f,\mathrm{max}} = 5$

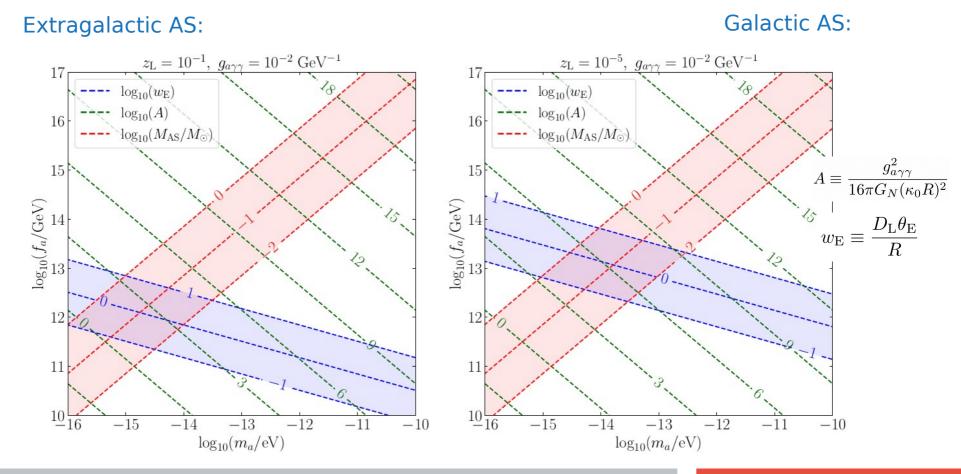
OMU-NTHU meeting 2025

P.Y. Tseng



Parameter regine

 Parameter region: gravity and axion-photon lensing are comparable:



OMU-NTHU meeting 2025

P.Y. Tseng

p.34

Sensitivities

Sensitivities:

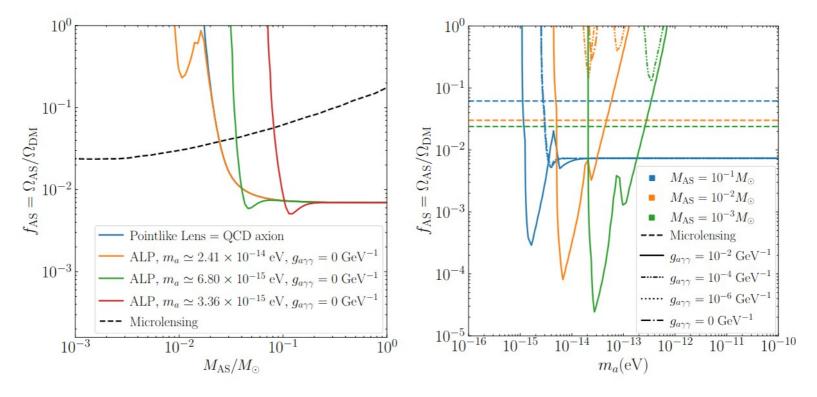


Figure 12: Constraint curves with $N_{\rm obs} = 10^4$ and $\Delta t_{\rm min} = 1 \,\mu s$ in the $f_{\rm AS}-M_{\rm AS}/M_{\odot}$ plane and $f_{\rm AS}-m_a$ plane, for the case of axion stars made of QCD axions and ALPs of fixed m_a (left panel) and ALPs (right panel).

J.T.Acuna, K.Y.Chou, P.Y.Tseng, arXiv:2501.07176

p.35

OMU-NTHU meeting 2025

P.Y. Tseng

Summary

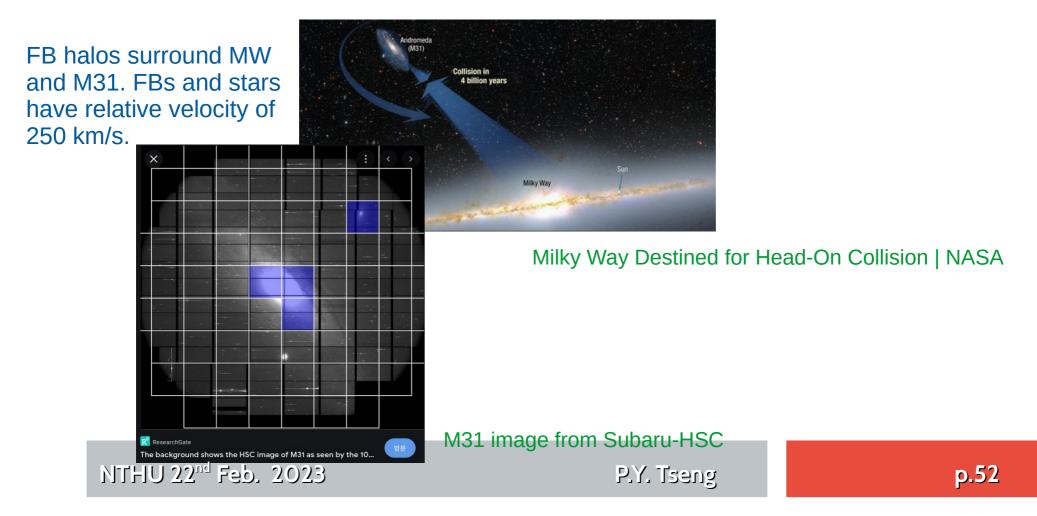
- We investigate the signatures from various DM candidates, including **PBH** and **axion star**.
- Light **PBH** features Hawking evaporation, generating gamma-ray or neutrino (cosmic-ray).
- **FOPT** provides a production mechanism for light **PBH**s and predicts Gravitational Waves.
- For axion star, we study the lensing effect induced by gravity and axion-photon couplings, considering signal of FRBs.

Thank you for your attention!

Back up

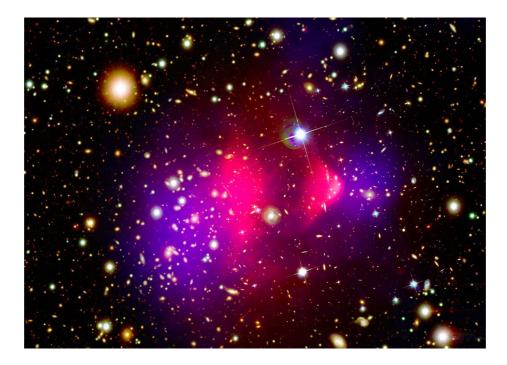
Microlensing

 Astrophysical Sky surveys are ideal for observing microlensing. Ex. Subaru-HSC (observing M31 for 7 hrs).



Indirect evidences of DM

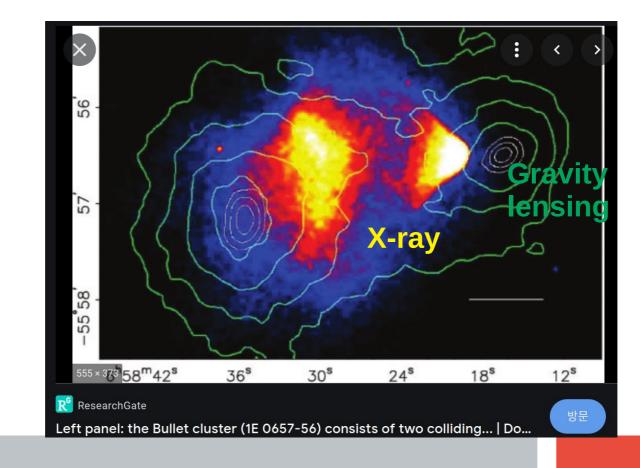
 Supporting evidence from Bullet cluster: picture of the two galaxies collided and pass through each other.



NASA: bullet cluster

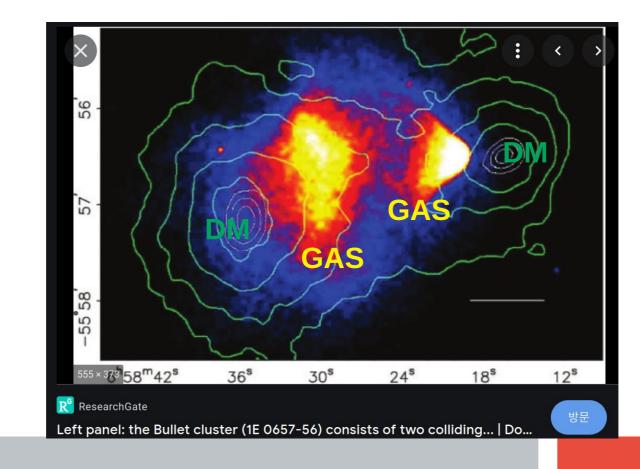
Indirect evidences of DM

 Supporting evidence from Bullet cluster: picture of the two galaxies collided and pass through each other.



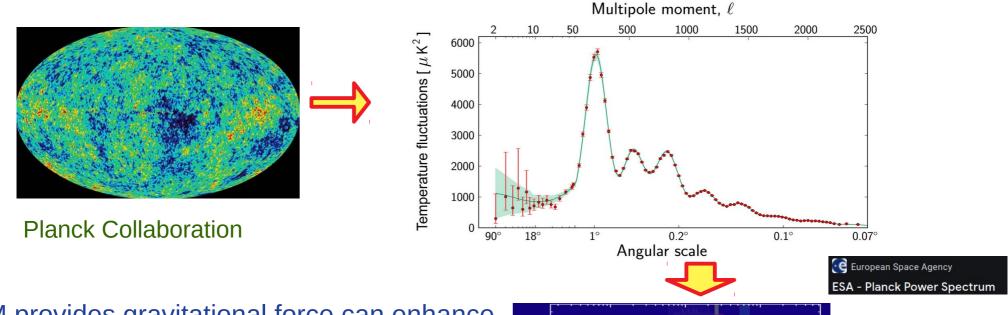
Indirect evidences of DM

 Supporting evidence from Bullet cluster: picture of the two galaxies collided and pass through each other.

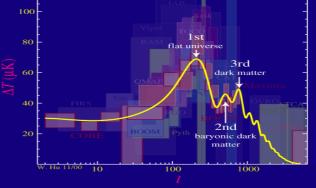


Indirect evidences of DM: CMB

• Expand the temperature by spherical harmonic functions:

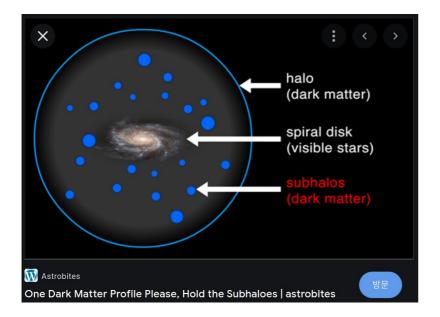


DM provides gravitational force can enhance the 1st and 3rd peaks, but weak the 2nd peak.



DM Search

- DM direct detection.
- We can detect more DM signal in summer than in winter.



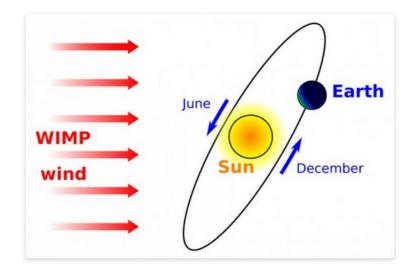


Illustration of the expected "WIMP wind" due to the motion of the Sun relative to the DM halo of the Milky Way. Figure from <u>arXiv:1209.3339</u>

- DM direct detection.
- Experimentalists put a detector underground shielding the background, mostly come from cosmic ray.

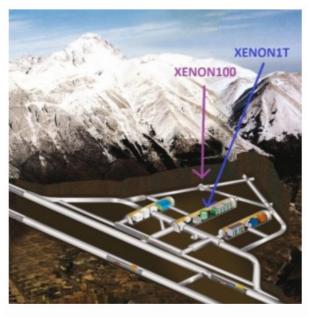
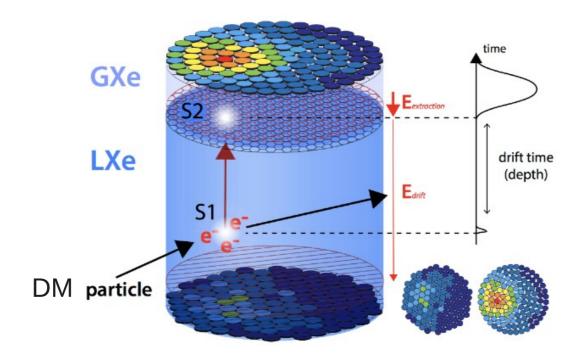


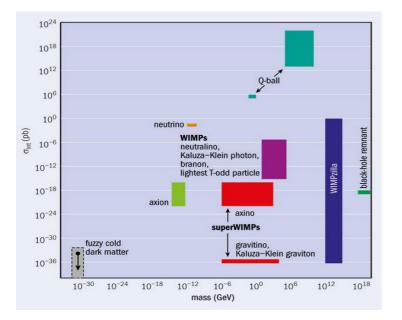
Figure 1: XENON100 and XENON1T Locations at LNGS

XENON1T detector:

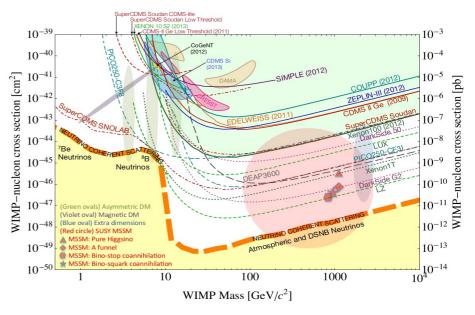


W Ni Group at UC San Diego - University of California San Diego [PDF] Status and Results of the XENON1T Dark Matter Search

- Supersymmetry theory predicts the WIMP (weakly interactive massive particle) around O(100) GeV to 1 TeV.
- It is constrained from direct detection searches.



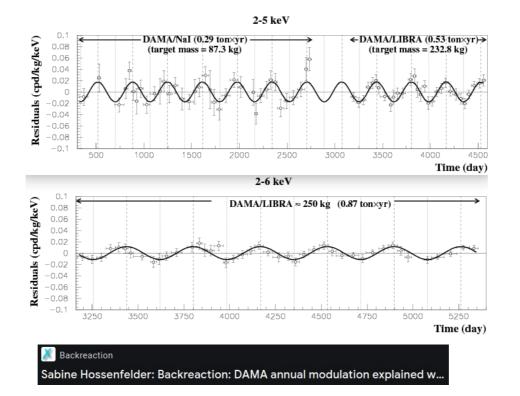




P. Cushman et. al.:1310.8327v2

DAMA/LIBRA (NaI) annul modulation signal.

But it can be explained without DM. (arXiv:1407.1052)



DAMA/LIBRA: arXiv:0804.2741