

WIMP-Nucleus Scattering: The Mean-Field Approach and Beyond

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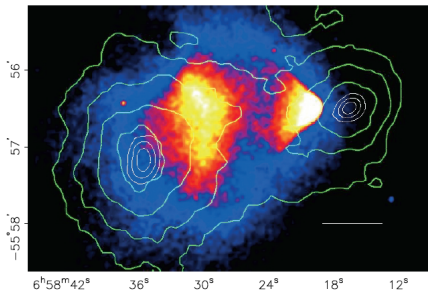


Figure 1: Mass distribution of the Bullet Cluster from gravitational lensing.
Image credit: 'A Direct Empirical Proof for the Existence of Dark Matter', Clowe, D. et al, ApJ 648 L109 (2006).

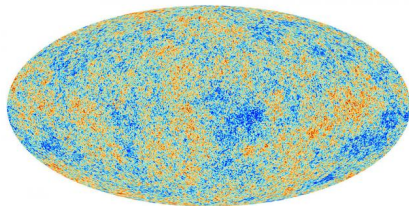
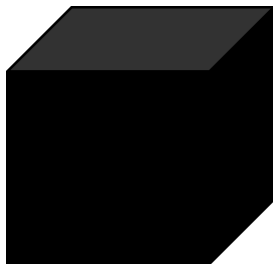


Figure 2: The cosmic microwave background.
Image credit: ESA and the Planck Collaboration.

Weakly Interacting Massive Particles

- WIMP: High-mass, spin-1/2 particle.
- Interactions below the electroweak scale.
- For this talk: we don't care about interaction specifics.



$$\frac{dR}{dE_R} = \frac{N_T \rho_\chi}{m_\chi} \int v f(\vec{v}) \frac{d\sigma_T}{dE_R}(\vec{v}) d^3v.$$

- Particle/nuclear physics input: differential cross-section.

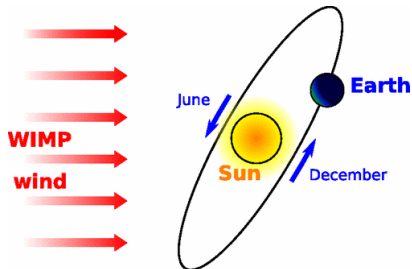
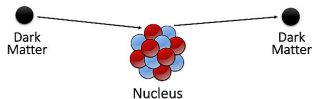


Figure 3: The 'WIMP wind' would generate an annual modulation in an interaction signal.

Image credit: Freese, K., Lisanti, M., Savage, C. (2013). 'Colloquium: Annual modulation of dark matter'. Reviews of Modern Physics, 85(4), 1561–1581.

$$\frac{d\sigma_T}{dE_R} = \frac{m_T}{2\pi v^2} \sum_{i,j} \sum_{N,N'=p,n} c_i^{(N)} c_j^{(N')} F_{ij}^{(N,N')}(q^2).$$

- Effective Field Theory Formalism [1]: $\mathcal{L} = \sum_i \bar{\chi} \mathcal{O}_i^X \chi \bar{N} \mathcal{O}_i^N N$.
- 15 possible WIMP-nucleon operators:
 $\mathcal{O}_1 = \mathbb{I}, \mathcal{O}_3 = i\vec{S}_N \cdot (\vec{q} \times \vec{v}^\perp), \dots$
- \Rightarrow 6 nuclear responses M, Δ, Φ'', \dots .
- Our focus: form factors

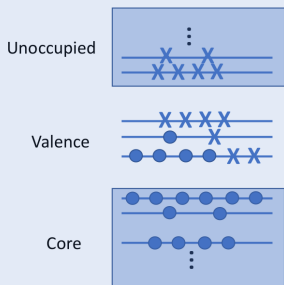
$$F_{X,Y}^{(N,N')} = \frac{4\pi}{2J_i+1} \sum_{J=0}^{2J_i} \langle J_i || X_J^{(N)} || J_i \rangle \langle J_i || Y_J^{(N')} || J_i \rangle.$$

[1] A. Liam Fitzpatrick et al. JCAP02(2013)004.

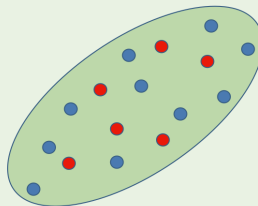
- Spherically symmetric, even-even nucleus $\Rightarrow J^\pi = 0^+$.
- Two relevant nuclear responses, M and Φ'' .
- $M_{JM}(q\vec{x}) = j_J(qx) Y_{JM}(\Omega_x)$.
Leading order: \mathbb{I} , counts nucleons.
- $\Phi''_{JM}(q\vec{x}) = i \left(\frac{\vec{\nabla}}{q} M_{JM}(q\vec{x}) \right) \cdot \left(\vec{\sigma}_N \times \frac{1}{q} \vec{\nabla} \right)$.
Leading order: $\sigma \cdot \ell$, nucleon spin-orbit coupling.

Nuclear Structure: Shell Model vs. Mean-Field

Shell model



Hartree-Fock-Bogoliubov



- Hartree-Fock: solve self-consistent Schrödinger equation.
- Skyrme force: phenomenological density-dependent force.
- Bogoliubov theory: add pairing interaction.
- Using code HFBTHO [2].

$$\begin{aligned}\hat{V}(1, 2) &= t_0(1 + x_0\hat{P}_\sigma)\hat{\delta}(x_1 - x_2) \\ &+ t_3(1 + x_3\hat{P}_\sigma)\rho^\alpha\hat{\delta}(x_1 - x_2) \\ &+ \dots\end{aligned}$$

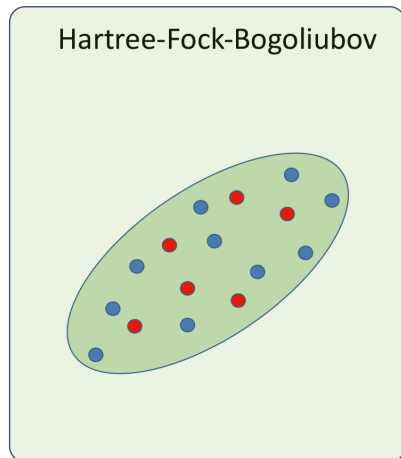
$$\hat{V}(1, 2) \rightarrow E[\rho] \rightarrow h_{ij}[\rho] = \frac{\delta E[\rho]}{\delta \rho_{ij}}.$$

$$\rightarrow h[\rho]\varphi_i = e_i\varphi_i.$$

[2]: Marević, P., Schunck, N., Ney, E. M., Navarro Pérez, R., Verriere, M., & O'Neal, J. (2022). Computer Physics Communications, 276, 108367.

Mean-Field: Correlations and Breaking Symmetries

- Localised \rightarrow break translational invariance.
- Deformed \rightarrow break rotational invariance.
- Pairing \rightarrow break particle number symmetry.

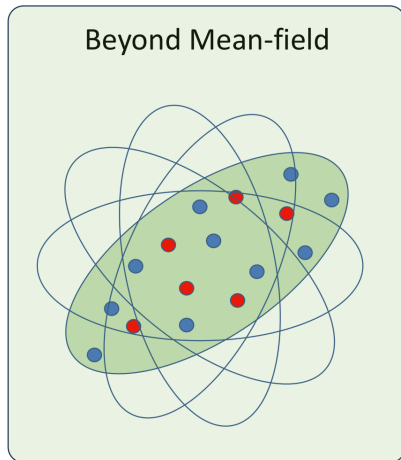


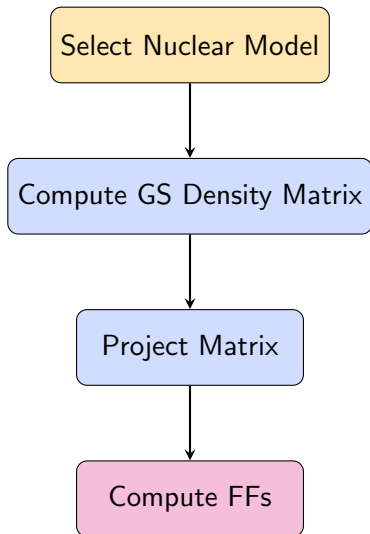
Restoring Symmetries

- Spin-dependent interactions couple WIMP and nuclear spin → need to project onto good angular momentum states.
- Pairing breaks particle number symmetry → need to project onto good particle number states.

$$P_N = \frac{1}{2\pi} \int_0^{2\pi} d\varphi e^{-i\varphi(\hat{N}-N)}.$$

$$|\psi\rangle = \sum_{N,Z} c_{N,Z} |\psi_{N,Z}\rangle.$$

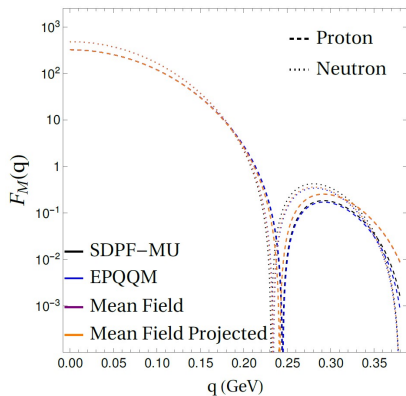




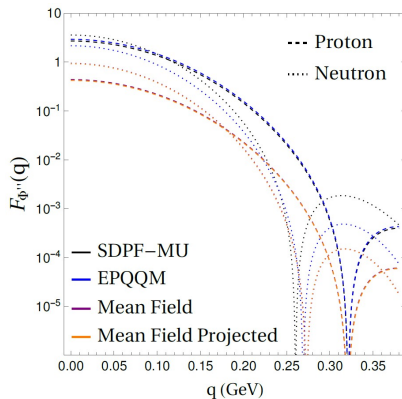
Results: ^{40}Ar Form Factors

$$F_{X,X}^{(N,N')}(q^2)$$

M - Contact Interaction

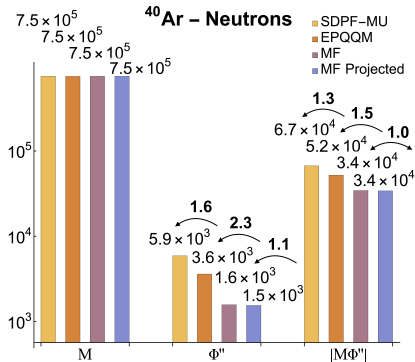
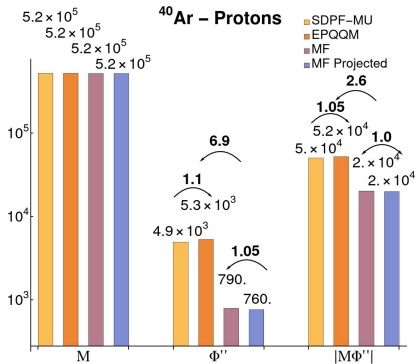


Φ'' - Spin-Orbit Interaction



Results: ^{40}Ar Integrated Form Factors

$$\int_0^{100 \text{ MeV}} \frac{q dq}{2} F_{X(Y)}^{(N,N)}(q^2)$$



- Need a mean-field approach to compliment existing work in WIMP-nucleus scattering.
- Demonstrated proof-of-principle of this approach in ^{40}Ar .
- Ongoing work: applying beyond mean-field techniques for odd and deformed nuclei.

Application of the Skyrme Hartree-Fock-Bogoliubov Theory to WIMP-Nucleus Interactions in ^{40}Ar

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(Dated: June 11, 2026)

WIMP scattering from ^{40}Ar is investigated using a self-consistent Skyrme Hartree-Fock-Bogoliubov (HFB) approach. Nuclear form factors relevant to dark matter direct detection are calculated from the resulting one-body density matrix elements and compared with shell-model predictions. Good agreement is found for the spin-independent response, while significant differences are observed for the spin-orbit response due to variations in single-particle occupancies. The effects of particle-number projection are shown to be small for ^{40}Ar . These results demonstrate the sensitivity of certain dark matter response channels to the underlying nuclear structure model and establish a framework for extending mean-field calculations to nuclei beyond the reach of large-scale shell-model studies.

arXiv: 2606.11668